

Wide-Bandwidth Near-Infrared Avalanche Photodiode Photoreceiver

Michael Krainak, Guangning Yang, Xiaoli Sun, Wei Lu
NASA Goddard Space Flight Center, MD

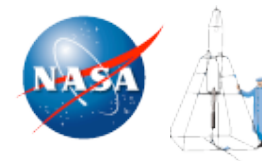
Xiaogang Bai, Ping Yuan, Paul McDonald, Joseph Boisvert,
Robyn Woo, Kam Wan, Rengarajan Sudharsanan
Boeing Spectrolab Sylmar, CA

Outline

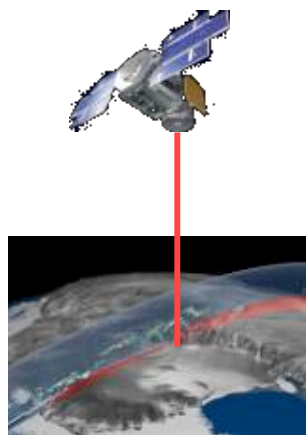


- NASA lidars
- Receiver Requirements
- Gen.1 photoreceiver design and performance
- Gen. 2 design and receiver development
 - I²E low excess noise APD design
- Summary

Space Borne Laser Altimeter Evolution ICESat/GLAS to LIST



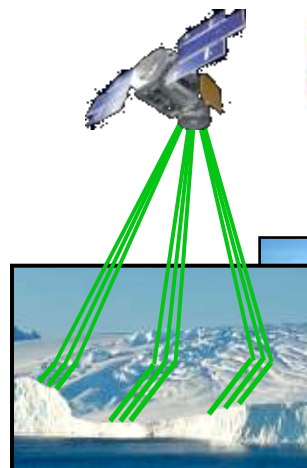
ICESat Single
Beam Profile



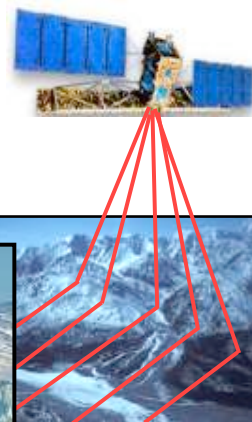
Ice Bridge
Airborne
Swath
Mapping



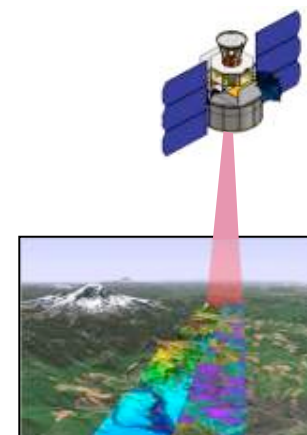
ICESat-2
Multi-Beam
Profiles



DESDynI
InSAR +
Multi-Beam
Profiles



LIST
Swath
Mapping



2005

2010

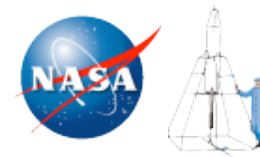
2015

2020

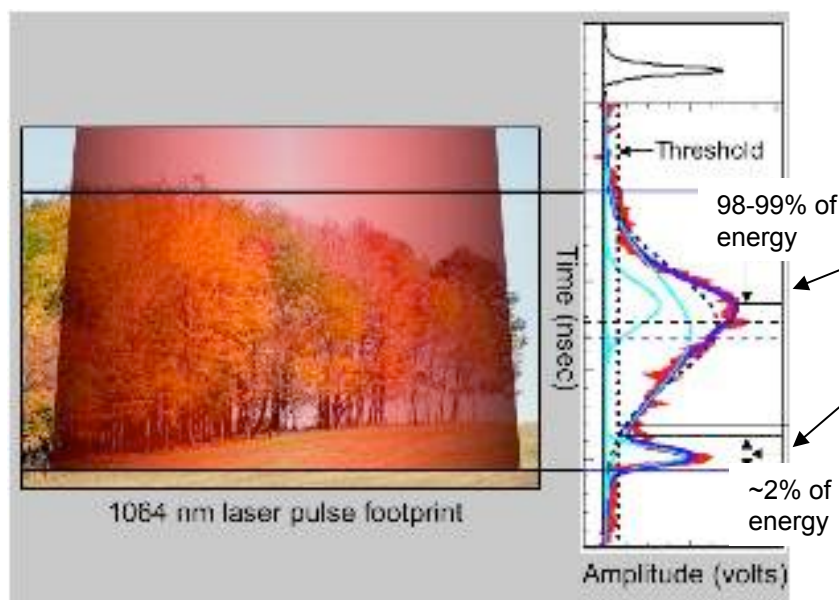
2025

*NRC Earth Science Decadal Survey
Missions*

Requirement to detect the weakest signal drives the lidar design (courtesy of J.Abshire)



A sample ICESat/GLAS Echo Waveform from tree(s)



Echo pulse waveform
(backscattered laser power vs time)

Echo pulse from Tree canopy

Echo pulse from ground under trees
Energy proportional to canopy opening fraction

For LIST, the opening fraction is ~2%

- GLAS acquires waveforms from vegetated terrain in a ~ 70 m diameter laser footprint

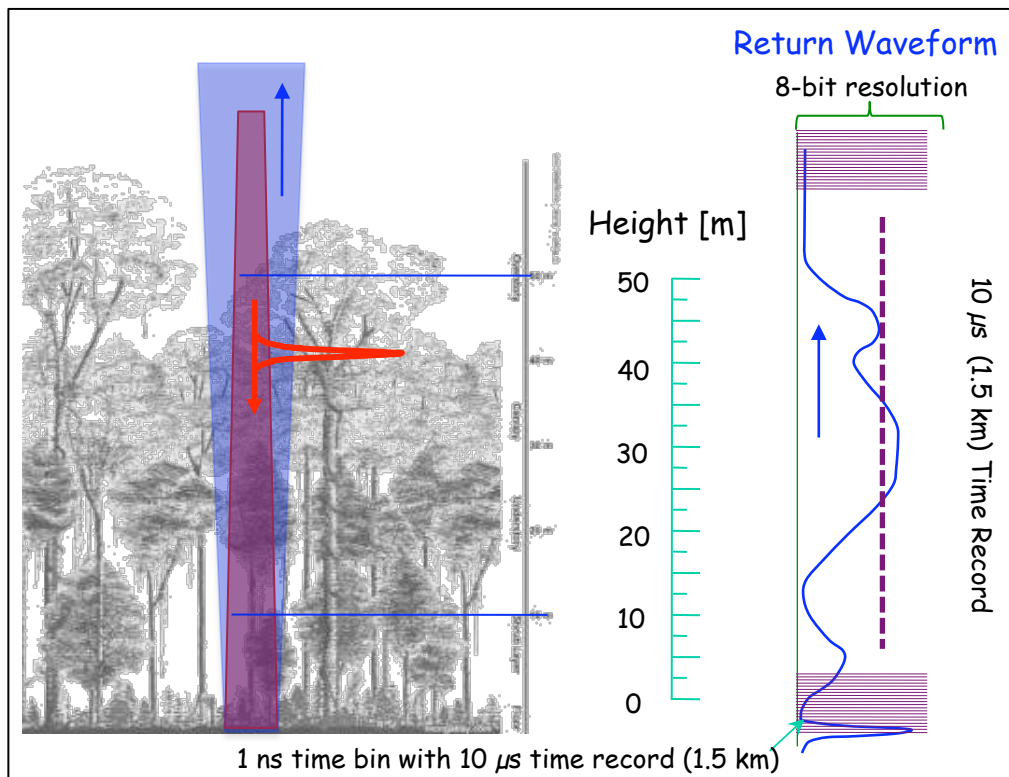
- Waveforms show height distribution of backscattered light reflected from canopy surfaces and underlying ground

☒ Energy in ground echo is ~2% of that from tree tops

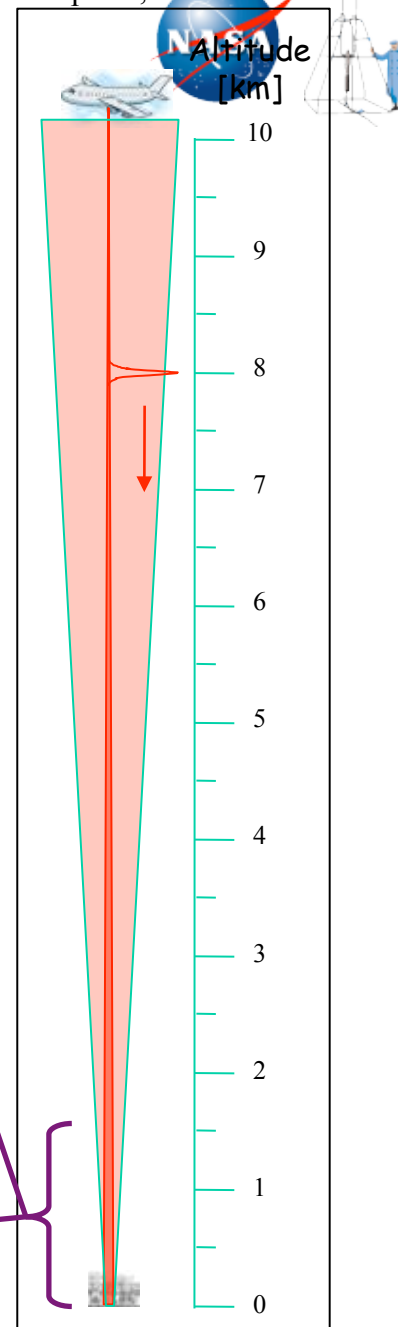
- Is weakest signal the altimeter receiver has to detect
- Energy scales with sub-canopy surface reflectivity
- Must reliably detect signal in presence of noise
 - (detected solar background rate)
- Need at least several detected photons/pixel

Lidar analog waveform processing

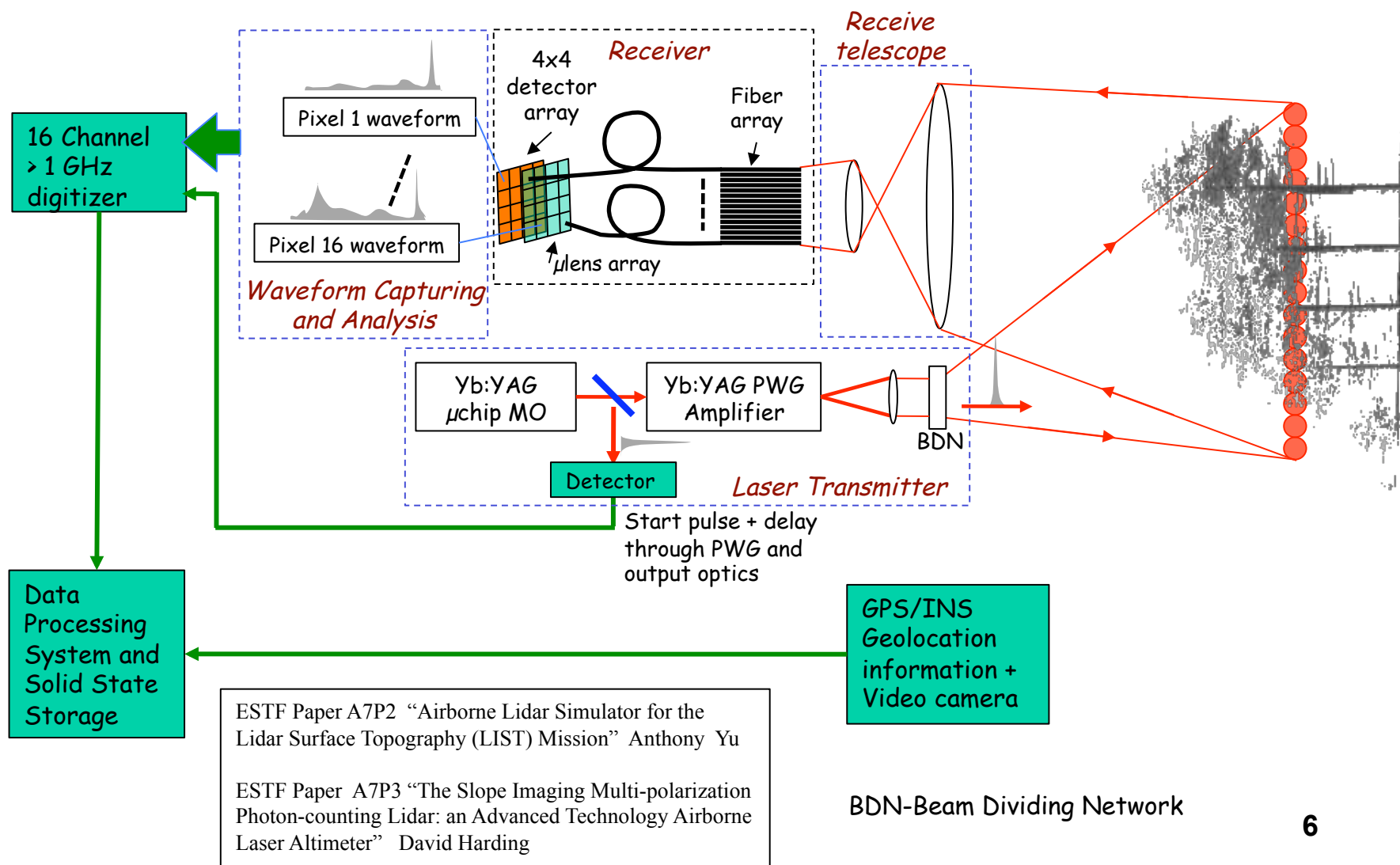
	Analog ICESat-2 approach	Swathmap IIP and LIST
Laser transmitter pulsewidth (ns)	6 ns	< 1 ns
Optical receiver/detector BW	160 MHz	1 GHz



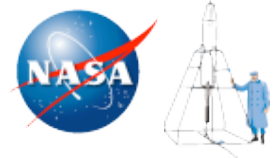
April 8, 2010



ESTO-IIP (Yu-PI) Airborne Instrument Concept

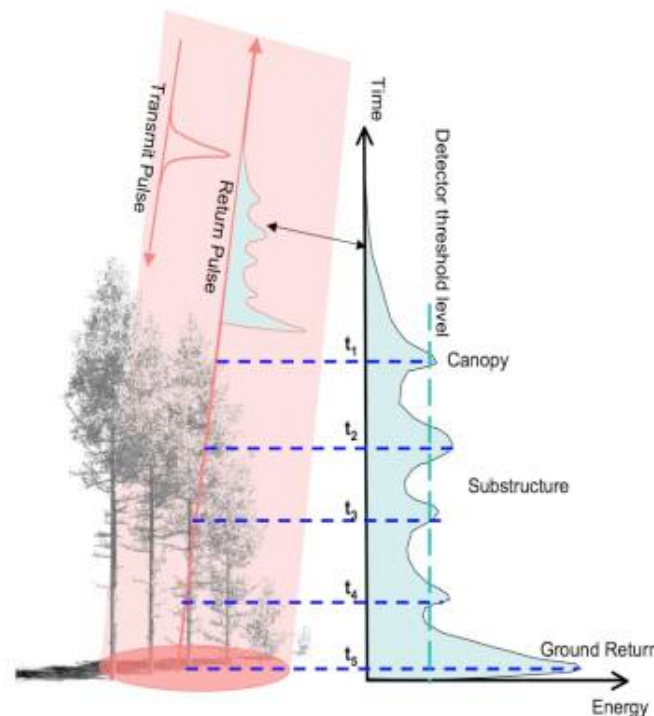


Receiver Requirements

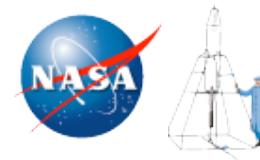


Generation 1: 1.06 μ m APD receivers with 200 μ m aperture, sensitivity $< 100 \text{ fW/Hz}^{1/2}$ @ a bandwidth of 140 MHz

Generation 2: 1.06 μ m APD receivers with sensitivity $< 300 \text{ fW/Hz}^{1/2}$ @ bandwidth of 1GHz



Lidar received pulse width variance



$$Var(T_R) = \frac{T_T^2}{\langle N \rangle} + \left(\frac{1}{\langle N \rangle} + \frac{1}{K_S} \right) \frac{4Var(\xi(\text{surface}))}{c^2 \cos^2 \phi_S} + \left(\frac{1}{\langle N \rangle} + \frac{1}{2K_S} \right) \frac{4z^2}{c^2 \cos^2 \phi_S} (\tan^4 \theta_T + \tan^2 \theta_T \tan^2 \phi_S)$$

$$K_S = \pi A_R \left(\frac{2 \tan \theta_T}{\lambda_0} \right)^2$$

“Target signatures for laser altimeters: an analysis” C. S. Gardner
APPLIED OPTICS /Vol. 21, No. 3 448 1 February 1982

$\langle N \rangle$ is the expected number of detected signal photons per received pulse

x_S is the surface profile variation (meters)

q_T is the laser beam divergence angle halfwidth at the $1/(e^2)$ point. (radians)

A_R is the area of the receiver aperture. (square meters)

λ_0 is the laser wavelength (meters)

ϕ_S is the surface slope angle (radians)

z is the altitude (meters)

c is the speed of light (meters/second)

T_R is the received pulse width (seconds)

T_T is the transmitted pulse width (seconds)

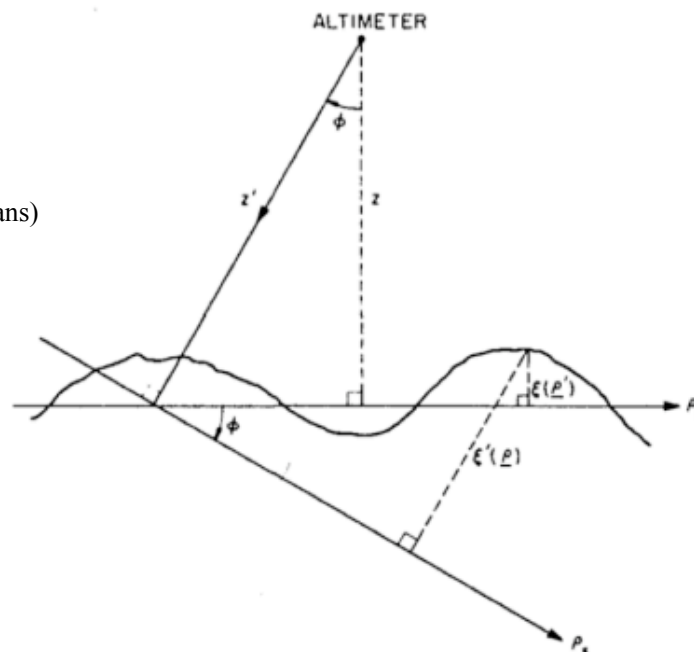


Fig. 2. Geometry of the laser altimeter and ground target for non-normal incidence.

Received pulse width variance

What conditions give an advantage to a shorter transmit pulse?

Case 1:
Airplane
10 km altitude
1 ns transmit pulse
5 m spot on the ground
0.5 m surface roughness

Surface slope (degrees)	Received pulse variance (ns)
0	0.19
1	0.19
2	0.20
3	0.21
4	0.22
5	0.24
6	0.26
7	0.28
8	0.30
9	0.33
10	0.36

Case 2:
Spacecraft
400 km altitude
1 ns transmit pulse
20 m spot on the ground
0.5 m surface roughness

Surface slope (degrees)	Received pulse variance (ns)
0	1.07
1	1.20
2	1.52
3	1.93
4	2.40
5	2.90
6	3.41
7	3.94
8	4.48
9	5.03
10	5.59

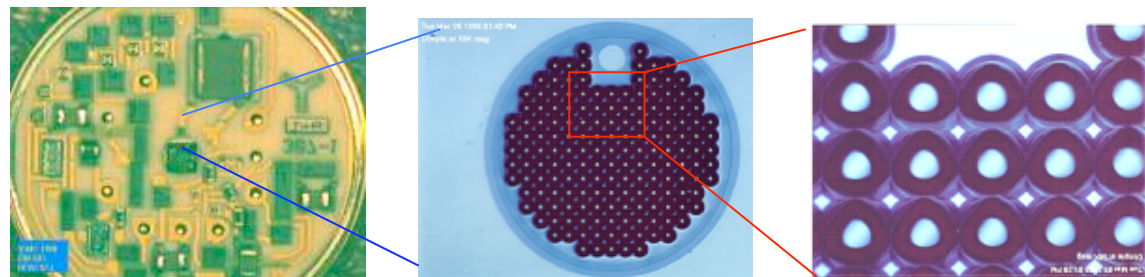
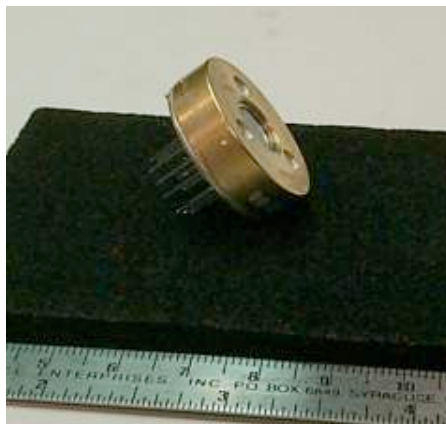
Case 3:
Spacecraft
400 km altitude
1 ns transmit pulse
20 m spot on the ground
0.1 m surface roughness

Surface slope (degrees)	Received pulse variance (ns)
0	0.22
1	0.58
2	1.09
3	1.62
4	2.16
5	2.70
6	3.24
7	3.79
8	4.35
9	4.92
10	5.49

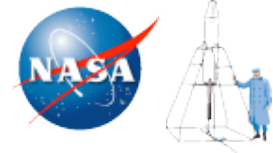
Near-Infrared Enhanced (Perkin-ELMER) Si APD optical receiver



- Near IR enhanced Si APD, preamp, and bias regulator hybrid in a 1" TO-8 package by PerkinElmer Optoelectronics Canada;
- Successful used in the laser altimeters in Clementine, NEAR, Mars Observer, ICESat, MESSENGER and LRO missions;
- The preamplifier was upgraded with a wide dynamic range and low distortion transimpedance amplifier during the ICESat development;



Near-IR Enhanced (Perkin-Elmer) Si APD optical receiver Performance



- **Quantum efficiency:** 40% (0.34 A/W) @1064 nm, 20°C (~50% at 65°C)
- **Active area diameter:** 700 μ m FWHM (800 μ m mask)
- **Gain:** 60-120
- **Excess noise factor** <3
- **Bulk dark current:** <50 pA
- **Bandwidth:** 140 MHz
- **Radiation damage:** No degradation after $10^{10}/\text{cm}^2$ protons (1-100 MeV)
- **Responsivity:** ~300kV/W
- **Preamplifier noise:** <1.5 pA/Hz^{1/2}
- **Noise equivalent power (NEP):** 30 – 40 fW/ Hz^{1/2}, dark
- **Linear dynamic range:** > 20dB in input optical signal
- **Power dissipation:** < 0.2 Watts
- **Operating temperature range:** 0 to 40 deg C
- **Lifetime:** >10 years in space (vacuum)

Noise Equivalent Power (NEP) Analysis



- APD + TIA Amp NEP
$$NEP = \frac{1}{R_{sp}} \left[2qI_d F + \frac{\alpha^2}{M^2} \right]^{1/2}$$

$$F = kM + \left(2 - \frac{1}{M}\right)(1 - k)$$

R_{sp} (=QE*q) is the APD unity gain responsivity

M is the APD optical gain

F is the APD excess noise factor

k is the ratio of the hole and electron ionization coefficients

α is the TIA noise current density

- Critical parameters to reduce NEP
 - Low ionization coefficient ratio, k
 - Low TIA noise α
 - High quantum efficiency
 - Optimize gain (M)

METHOD

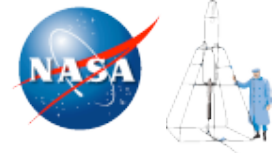
Engineer the material (I2E)

Electronics design

Choice of absorber material

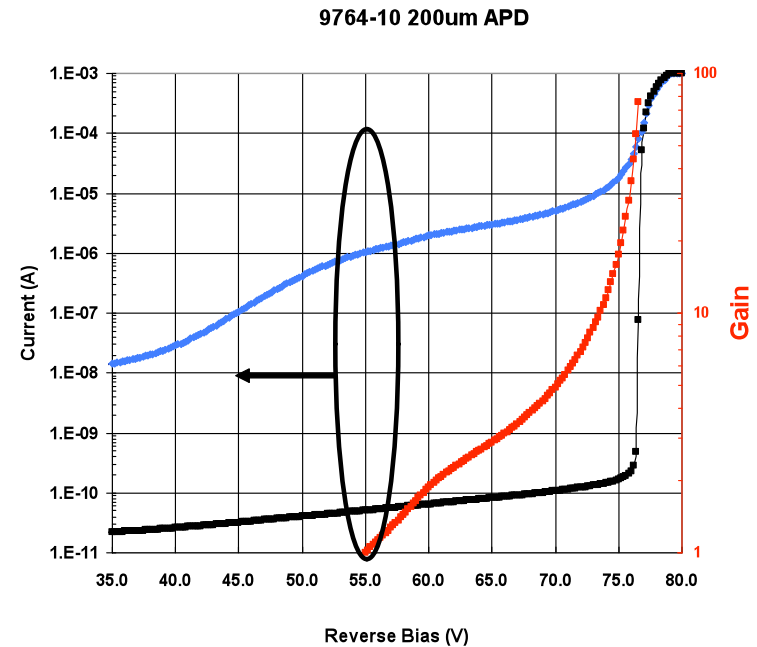
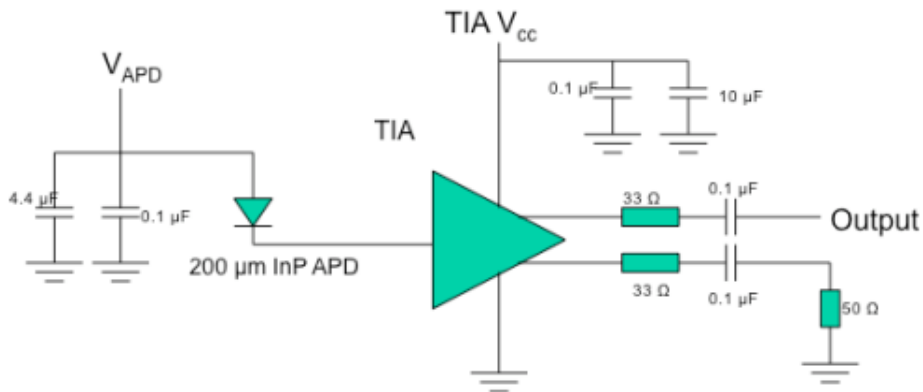
Engineer the material (I2E)

Gen. 1 Photoreceiver Design



- 200 μm InP APD – “off-the-shelf” material
- Low noise TIA , SA5211 1.8 $\text{pA}/\text{Hz}^{1/2}$
- An integrated TEC cooler and a AD590 temperature sensor chip

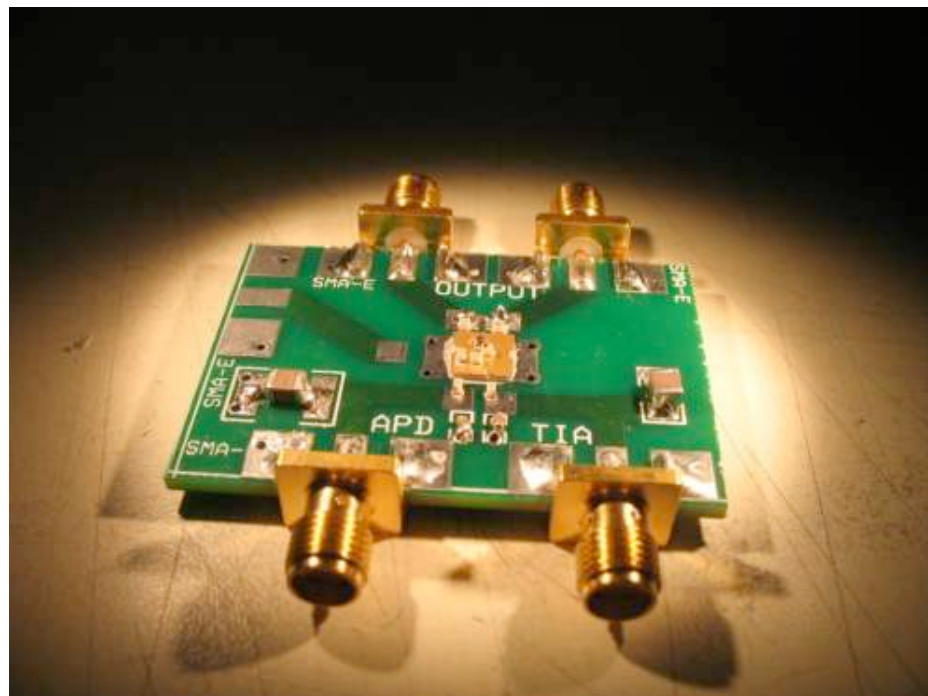
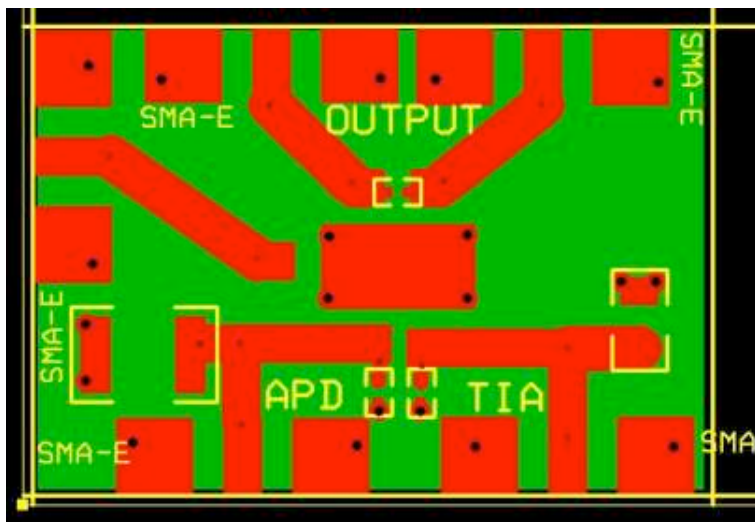
Receiver I-V data



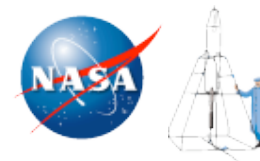
Photoreceiver



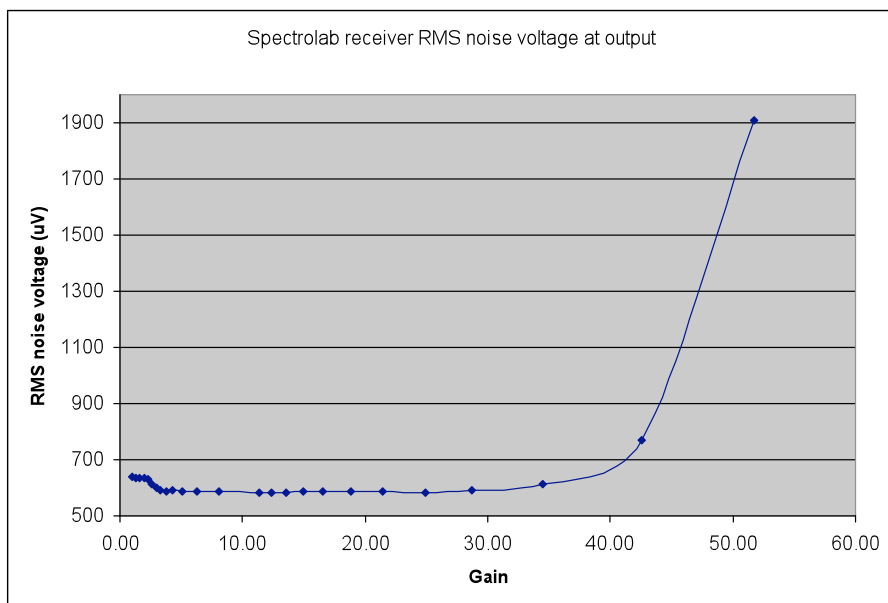
Photoreceiver = APD + Transimpedance amplifier



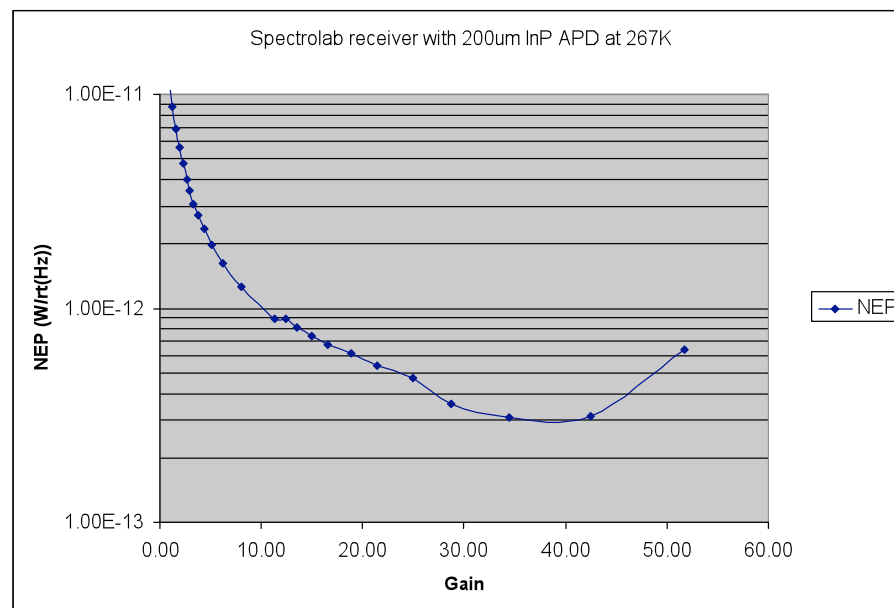
Gen. 1 Photoreceiver NEP Data



RMS Voltage Data



NEP Data

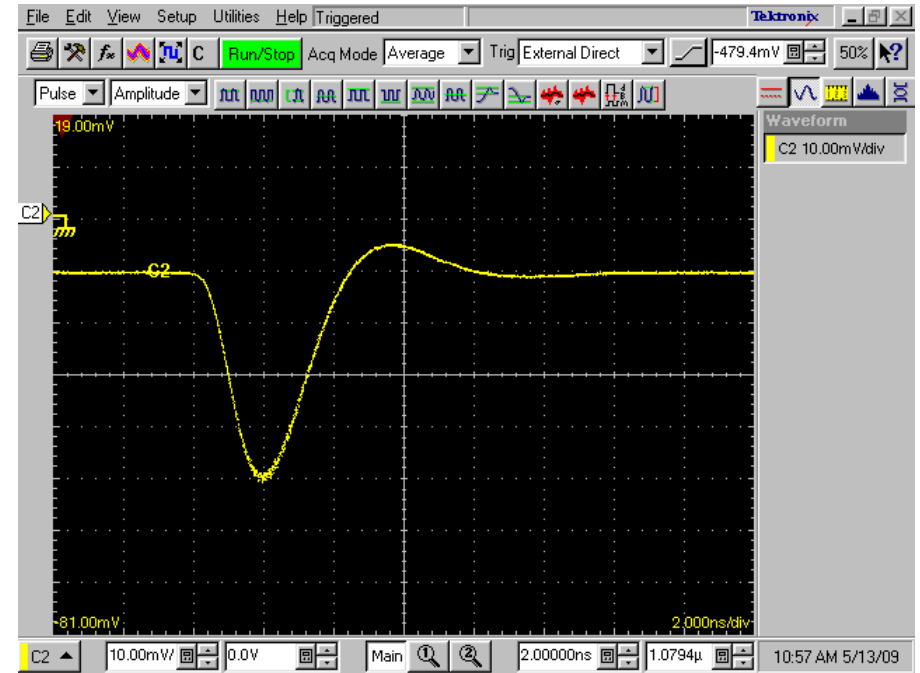
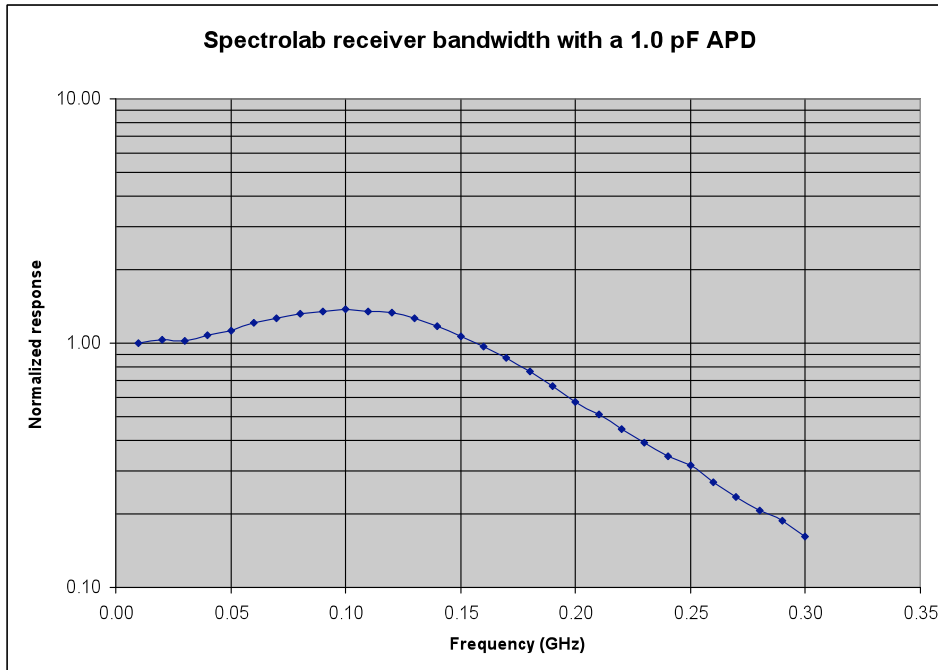


- NEP < 300 fw/Hz^{1/2} was achieved

Gen. 1 Receiver Bandwidth



Receiver response to a 100ps 1.06 μ m laser pulse.



- Achieved bandwidth of 180 MHz

Performance metrics



Parameter	Performance	Unit	Approach
Wavelength	1.06	μm	Device
NEP	300	$\text{fW}/\sqrt{\text{Hz}}$	Device, circuit
Bandwidth	>1	GHz	Device, circuit
Quantum Efficiency	>75	%	Device
k_{eff}	<0.15	-	Device
Fill Factor	>70	%	Microlens
Number of channels	16		APD array

Gen-2 Receiver Sensitivity Improvements



Goal: $300 \text{ fW/Hz}^{1/2}$ @ bandwidth of 1GHz

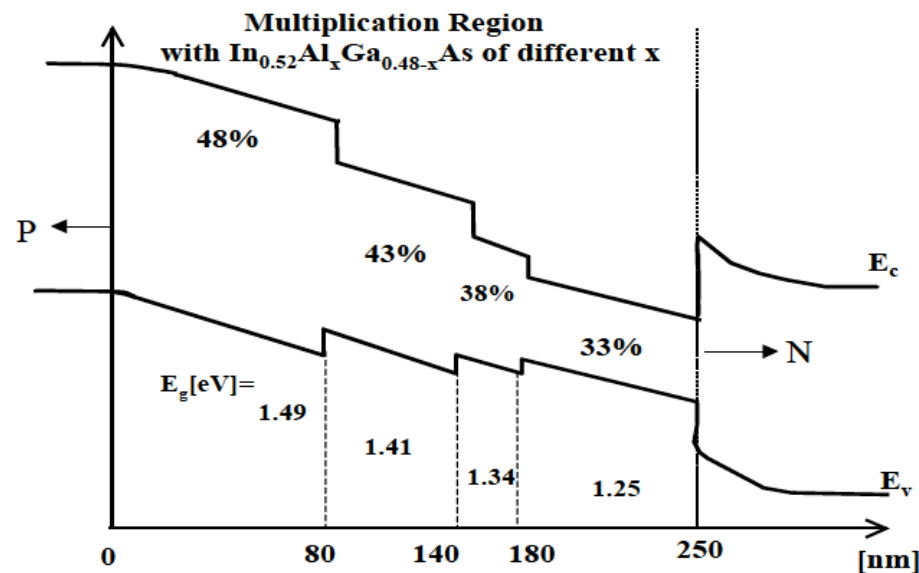
- Quantum efficiency
 - Gen. 1 InP APD has 64% quantum efficiency
 - 75% QE will reduce NEP by 17%.
- Low noise TIAs
 - Select best low noise TIAs in die form with less than $6 \text{ pA/Hz}^{1/2}$ input referred noise current.
- Reduce excess noise in APD
 - InAlAs has a k value ~ 0.22
 - Engineer a material (I^2E APD design) with reduced $k_{\text{eff}} \leq 0.15$

I²E APD Fundamentals



- APDs have high internal gain and associate excess noise
- k factor is a material parameter for bulk material

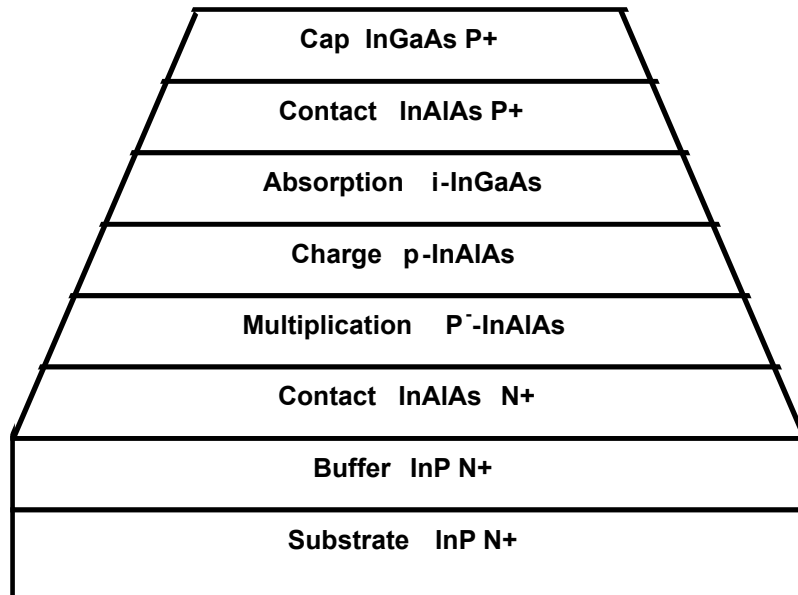
I²E= Impact Ionization Engineering



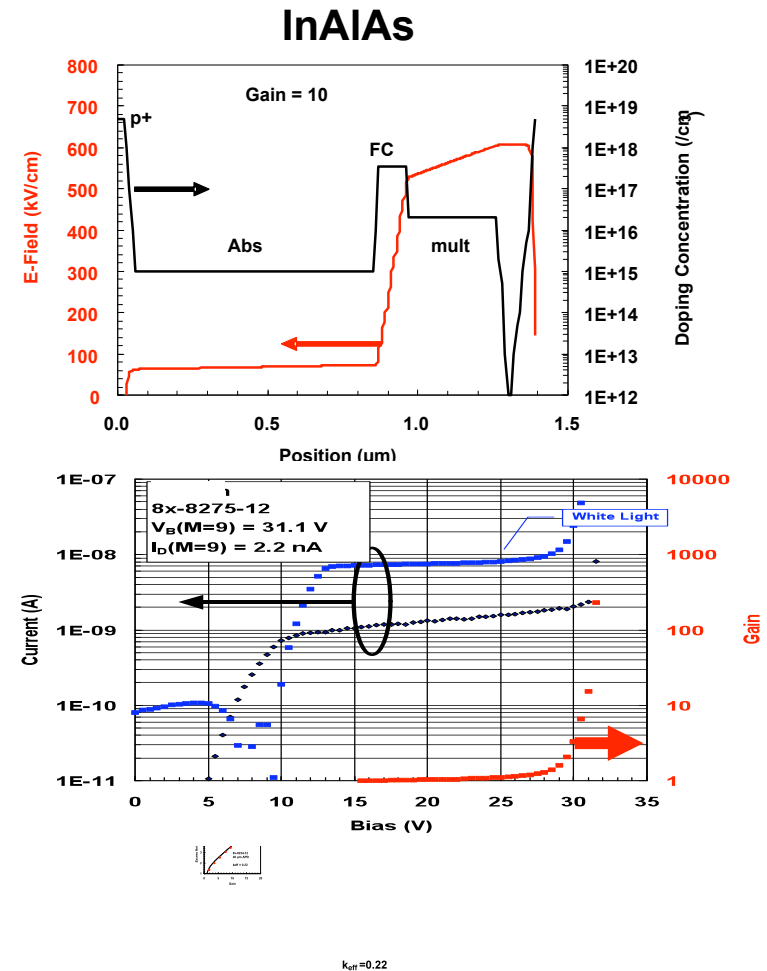
I²E is an approach to combine materials with different impact ionization threshold energies in the multiplication region. In the I²E structure, the avalanche events are more deterministic which result in a low effective k-factor (i.e lower excess noise). A lower bandgap material typically has a lower ionization energy threshold.

Graph from S. Wang, et. al., IEEE Photonics Technology Letters, Vol.14, No. 12, pg1722, 2002

Spectrolab InAlAs APD



InAlAs APD shows $k_{\text{eff}} = 0.22$



Spectrolab I²E APD Design



p+ InGaAs Cap layer, 50nm

p+ InAlAs, 300nm

i-InGaAlAs Absorber
1200nm
E_g~1.05 eV

p+, InAlAs Charge layer

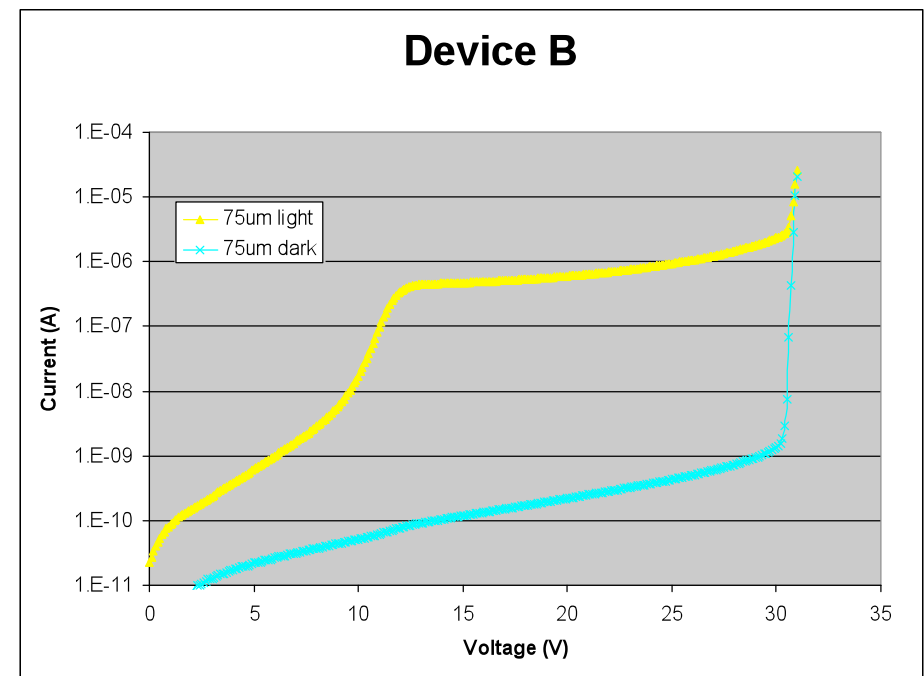
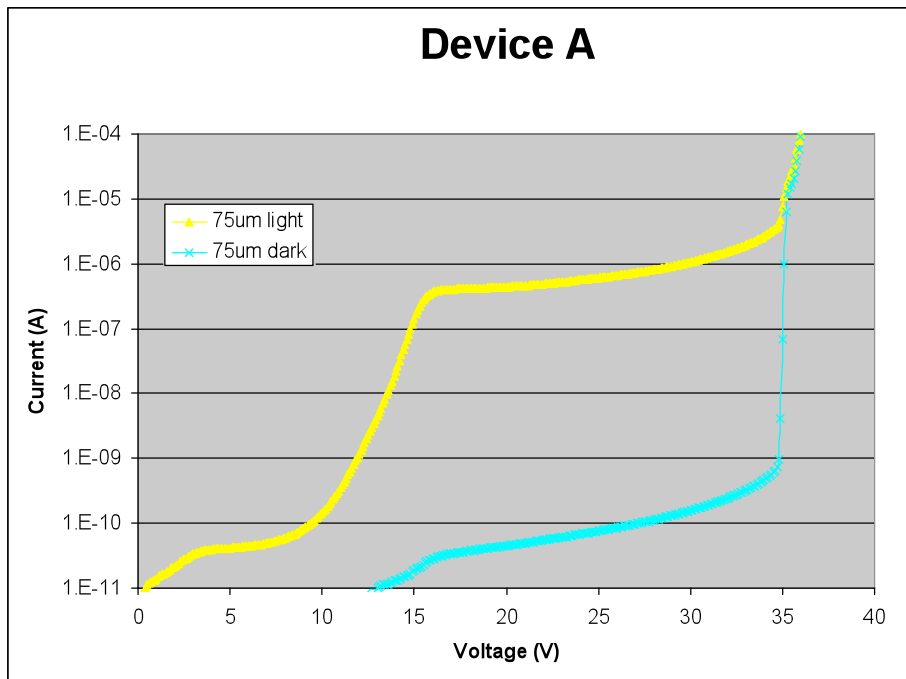
I²E Multiplier

n+ InAlAs Buffer

n+ InP Substrate

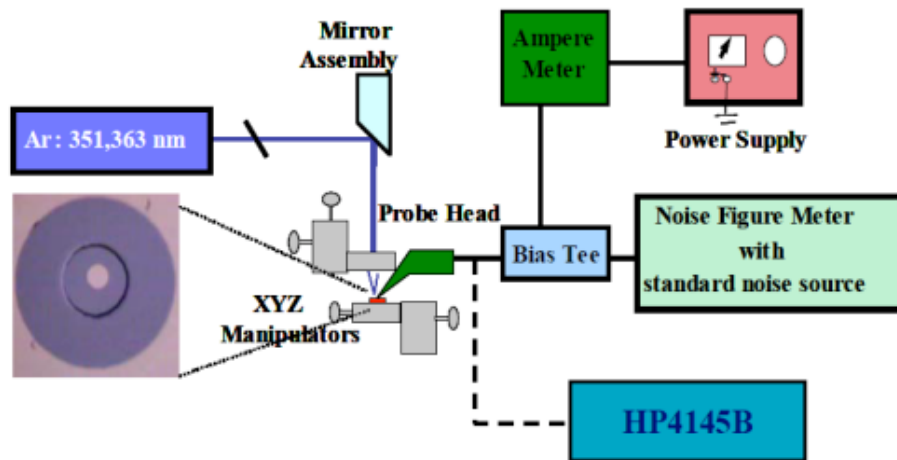
- InGaAlAs layer with bandgap of 1.2 eV is used as a multiplier

I²E Device I-V Data



- Show very low dark current before breakdown.

Excess Noise Measurement



$$S = 2eI_{\text{unity}} M^2 F(M) R(\omega)$$

Testing procedure:

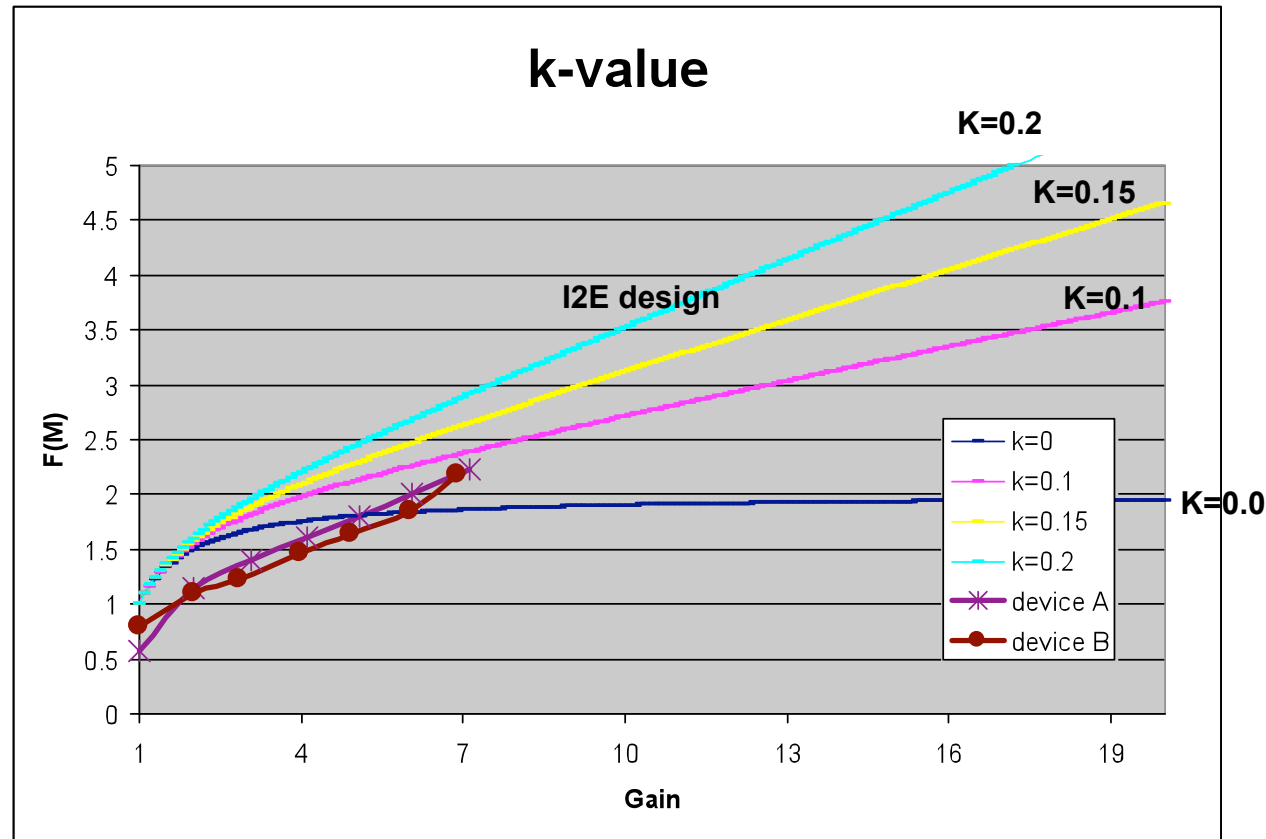
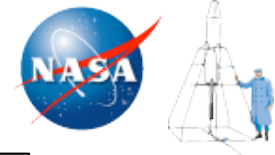
(a) At unity gain, measure S vs. I_{unity} to fit the $2eR(\omega)$.

(b) Measure S vs. M to get $F(M)$.

UV laser is absorbed near the surface p+ contact layer.
Electrons are diffused into the multiplier, thus pure electron injection is realized.

* Setup graph is from Dr. Shuling Wang's Ph. D. dissertation(2002).

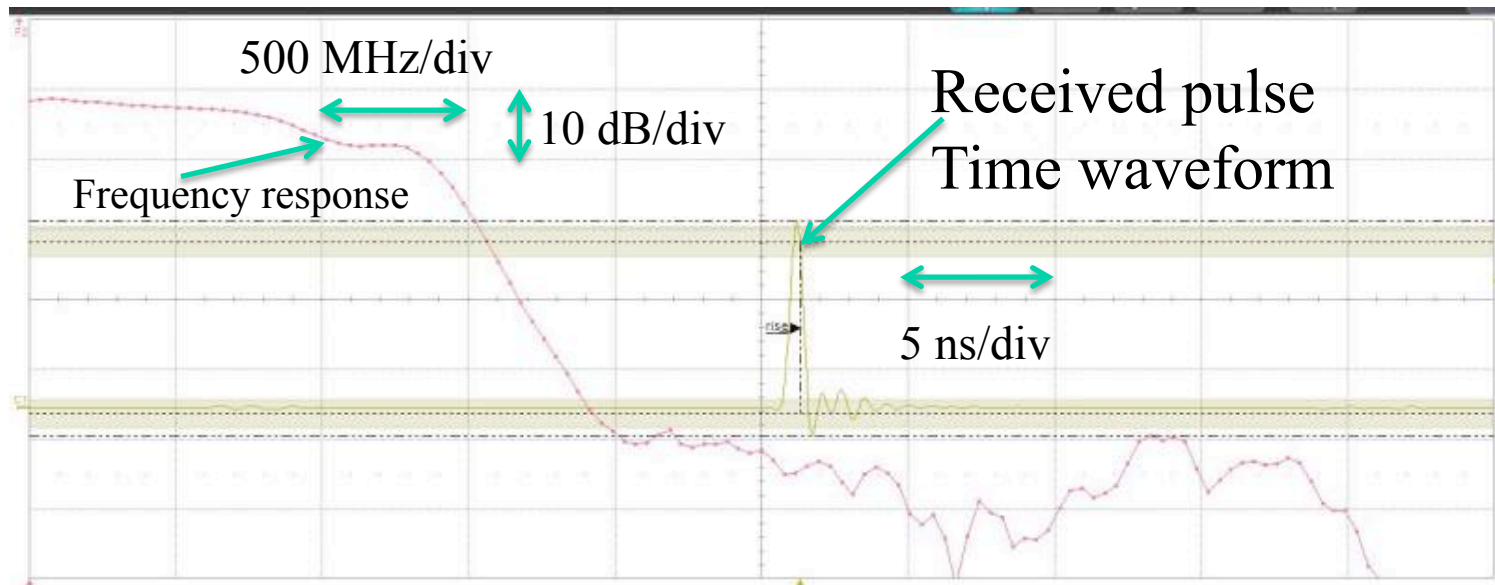
Excess Noise Results



- Both device A and B show k value less than 0.1
- $k \leq 0.15$ is feasible at high gain (15~30) for future I²E

I2E APD (1st run) Pulse/Frequency Response

Waveforms are captured with average of 1024

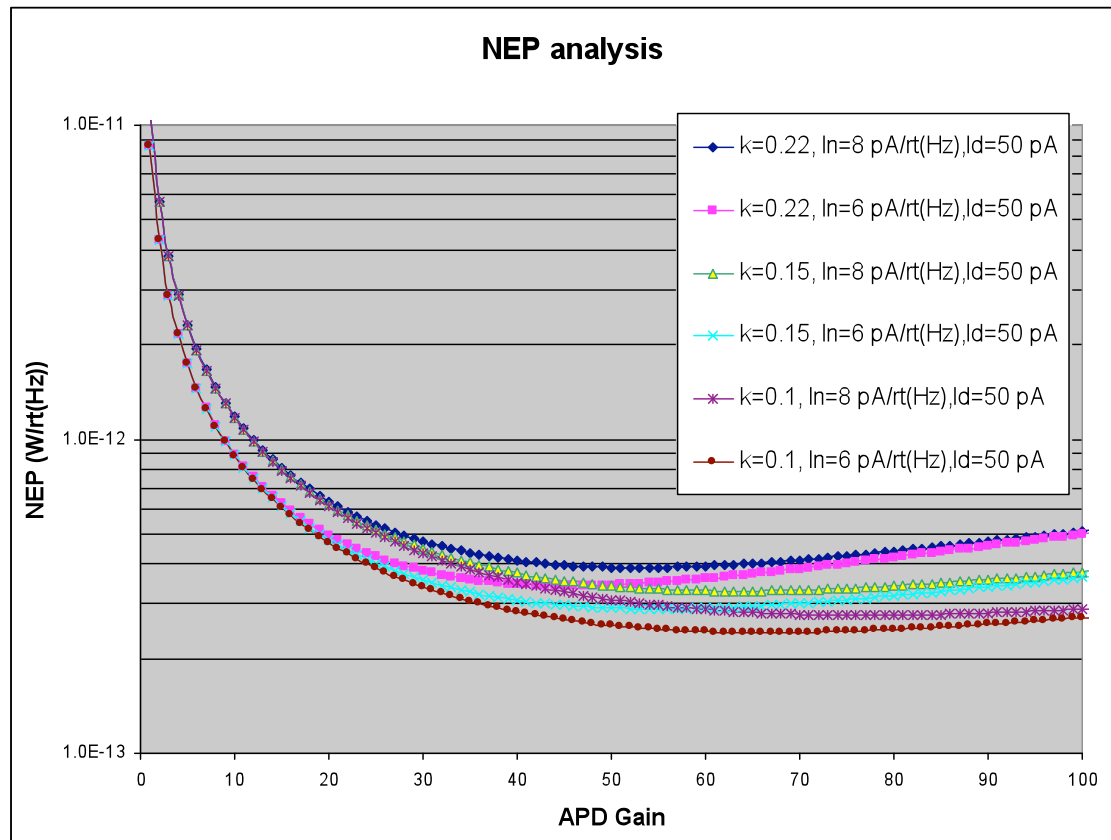


The detector BW is at least 1.5 GHz

Impulse response
Source=PicoQuant Pulse Laser
Laser Input Pulse = 98ps FWHM
Gaussian

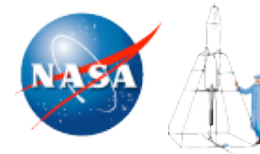
BW is limited by the pre-amp
ZFL1000 (1 GHz) and HP8447 (1.3 GHz)
Bias is set to 34.4V
APD Current is measured 0.435mA

Gen. 2 Photoreceiver – NEP Analysis



- NEP less than $300 \text{ fW/Hz}^{1/2}$ over 1GHz bandwidth can be achieved using I²E devices

Recent results – June 2010

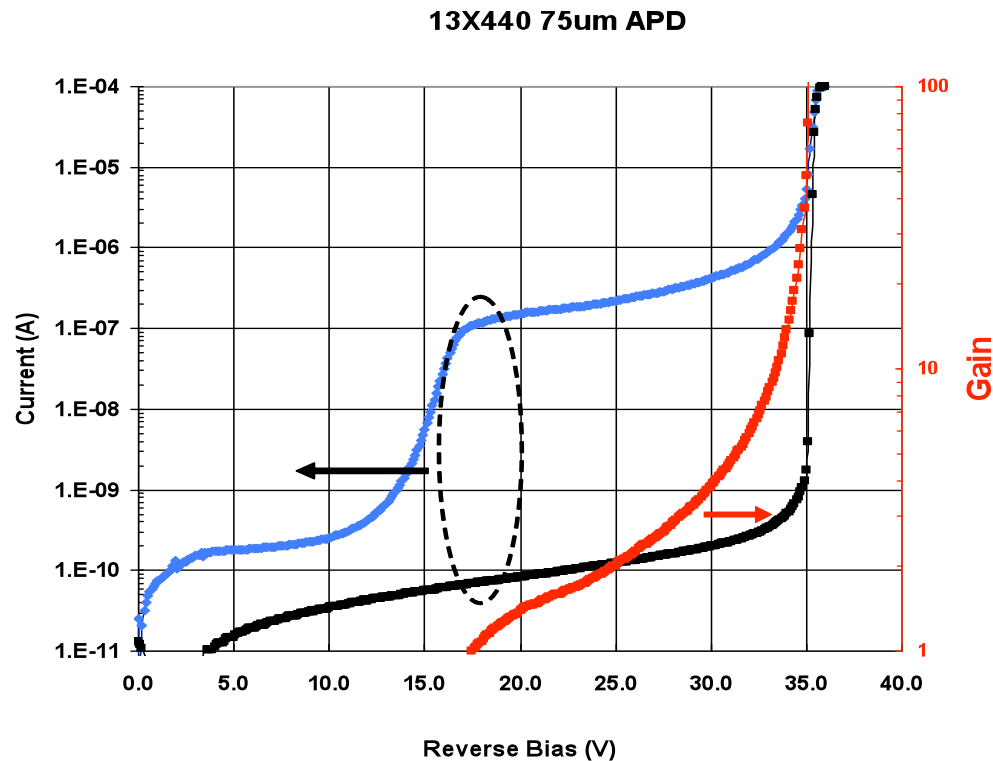


- From the first run (late 2009), the I²E APDs showed low excess noise of $k \sim 0.1$ but device gain was limited < 7 .
- In 2010, Spectrolab re-grew the same I²E APD structure in a new reactor and optimized the growth conditions to increase the gain over > 10 . **A gain over 50 was achieved** and we assume the same low excess noise since the structure is same. We will re-measure the k-value soon.
- We also designed and built 1x16 APD array and investigated the uniformity. We observed very good uniformity.
- Designed and developed 2nd Gen I²E APDs to achieve low capacitance. We observed 50% reduction in capacitance.

I²E low excess noise APDs



- Room temperature measurement.
- 200 micron devices
- The photocurrent was measured with APDs illuminated with 1.06 μm laser.



- Improved the gain in our I²E low excess noise APDs

2nd Gen I²E APDs

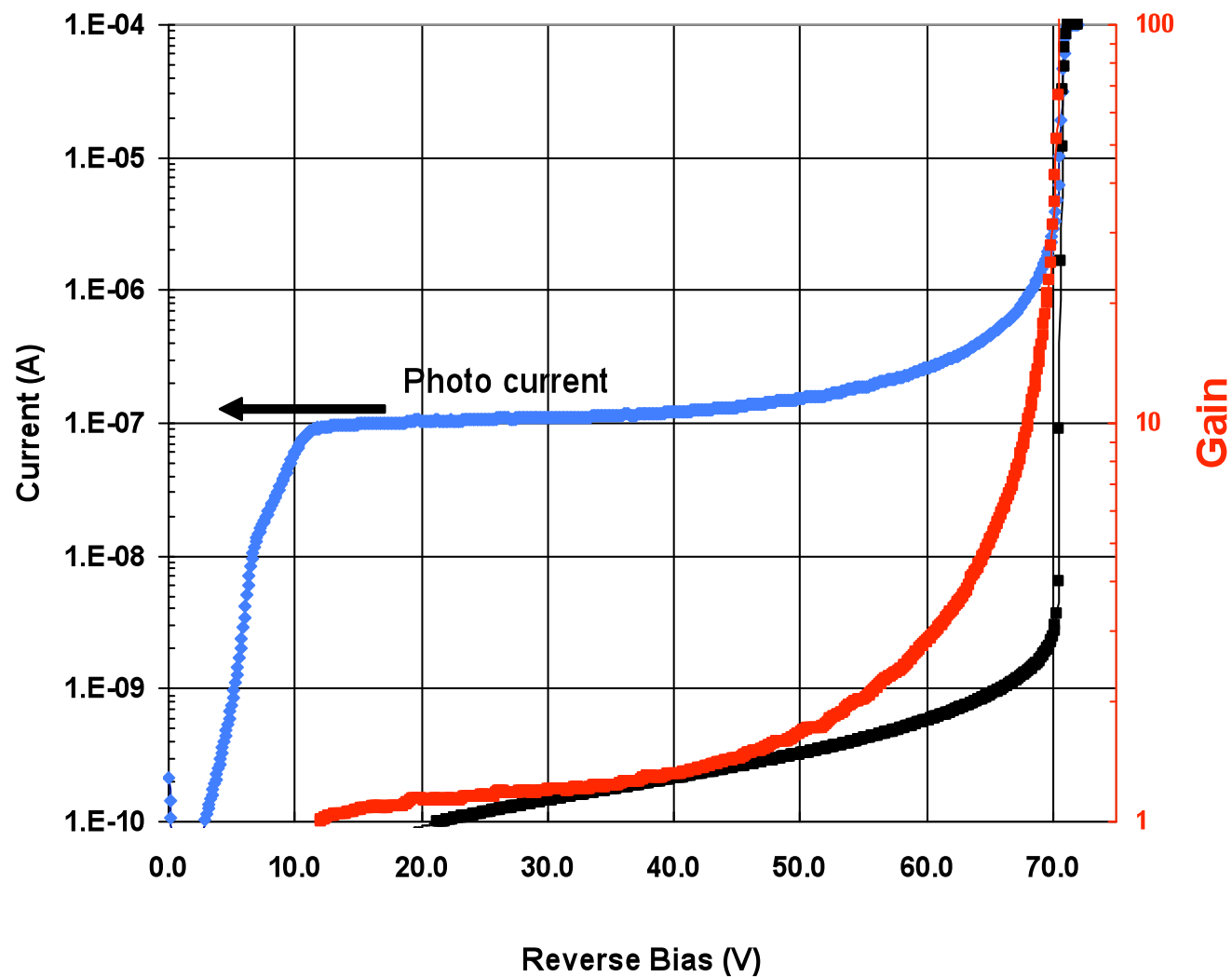


- At working gain, the capacitance of 1st generation I²E APDs is 0.62 pF 75 μ m devices
- The capacitance of 2nd Gen I²E APDs is 0.3 pF for 75 μ m devices, a 50% reduction in capacitance
- Low capacitance will reduce receiver noise

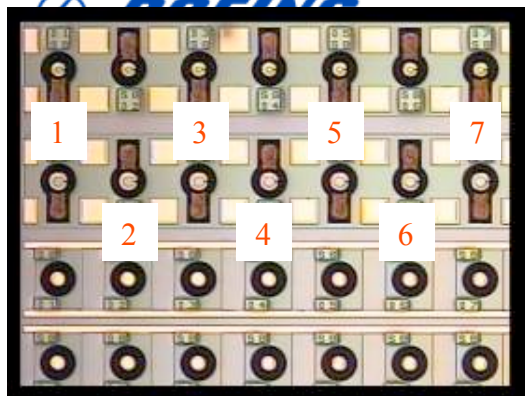
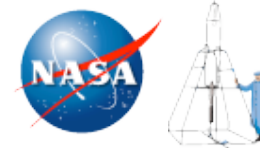
2nd Gen I²E APD IV



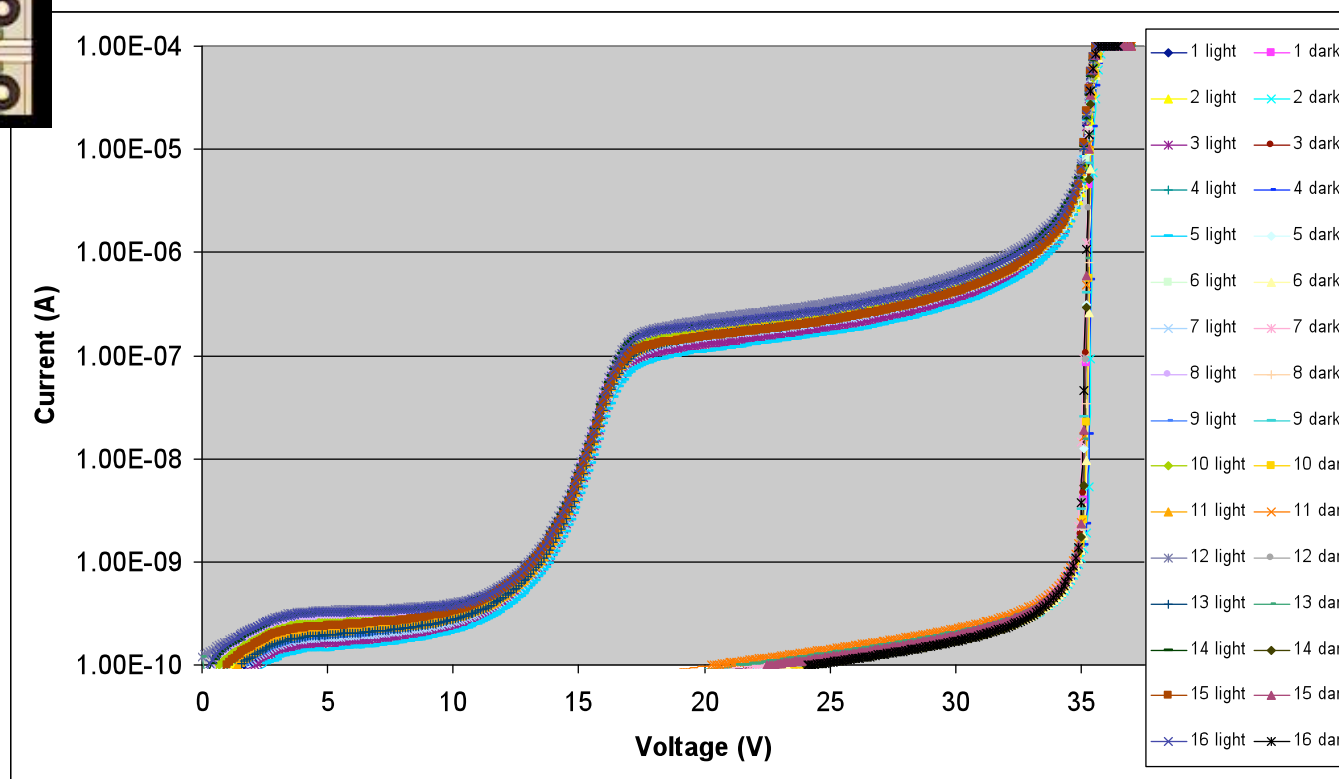
13X447 75um APD



1X16 I²E APD array



- 1x16 I²E APD array with a pitch of 250 microns.
- Photo current is under 1.06 μm laser illumination.
- Breakdown voltages are in a range of 0.2 volt.



- Demonstrated good array uniformity

Summary



- Full-waveform processing lidar receivers provide surface topography, tree canopy and atmospheric data.
- Short pulsed lasers (1 ns) and high-bandwidth (1 GHz) photoreceivers provide improved data:
 - ◆ for airborne lidars
 - ◆ for spaceborne lidars over regions with low surface slope and low surface profile variation
- Demonstrated $NEP < 300 \text{ fW/Hz}^{1/2}$ photoreceiver using InP APD
- Developed InAlAs based I²E APDs with > 1 GHz bandwidth.
- Demonstrated low excess noise APDs, $k_{\text{eff}} < 0.1$
- Demonstrated gain > 50 .
- Developing Gen. 2 receiver with I²E APD devices to achieve NEP less than $300 \text{ fW/Hz}^{1/2}$ over 1 GHz